THE EUROPEAN PHYSICAL JOURNAL D EDP Sciences © Società Italiana di Fisica Springer-Verlag 2001

# Extended analysis of the six-times ionized xenon, Xe VII

M. Gallardo<sup>1,a</sup>, F. Bredice<sup>1,b</sup>, M. Raineri<sup>1,c</sup>, A.G. Trigueiros<sup>2</sup>, and J.G. Reyna Almandos<sup>1,d</sup>

<sup>1</sup> Centro de Investigaciones Opticas (CIOp), Casilla de Correo 124, B1902WAB La Plata, Argentina

<sup>2</sup> Instituto de Física, Universidade Federal Fluminense, 24210-310 Niterói, Río de Janeiro, Brasil

Received 11 December 2000

**Abstract.** The spectrum of six-times ionized xenon, Xe VII, has been studied in the 270-6500 Å using a pulsed discharge. 110 new classified lines are reported. Ten levels belonging to the odd parity configurations and four belonging to the even configurations have been determined. To obtain the energy parameters Hartree-Fock relativistic calculations were used. Least-squares parametric calculation has been carried out to study the fit between experimental and theoretical values.

**PACS.** 32.30.Jc Visible and ultraviolet spectra – 52.80.Yr Discharges for spectral sources (including inductively coupled plasma)

### 1 Introduction

Cadmium-like xenon, Xe VII, has the ground configuration  $5s^2$ . Previous works were carried out by different groups of researcher [1–6] who used various experimental methods such as beam-foil spectroscopy, spark-light sources, laser-produced plasma, theta-pinches, etc. A study of the 5p5d configuration by Cavalcanti *et al.* was in the last time reported [7]. Larsson *et al.* [8] obtained energy levels of the 5s6s, 5s5d and  $5p^2$ , and 5s4f configurations by collision-based spectroscopy. Using the same kind of excitation, Wang *et al.* [9] extended the Larsson's work to include the 5s(6p + 7p) and 5s(6d + 7s) configurations. New results for ions in the Cd-like isoelectronic sequence [10–14] were recently published.

In this work we report additional experimental results for the Xe VII spectrum in the VUV region. In these results we include 14 new energy levels and 110 new transitions. The accuracy of the confirmed values of the known energy levels reported in references [8,9] were also improved. This study is a continuation of our investigations of different xenon ion spectra using gases pulsed electrical discharges ([15–19], and references therein). Hartree-Fock calculations including relativistic corrections and interpolation through the Cd I isoelectronic sequence were used to predict energy level values and transitions. It is expected that the spectroscopy of highly ionized xenon will be of interest for future diagnostics in the International Tokamak for Experimental Research, ITER ([7], and references therein).

### 2 Experiment

A light source specially built at CIOp to study highly ionized noble gases, based on the discharge tube reported by Gallardo et al. [19], was used in this work. It is a 110 cm long Pyrex tube with an inner diameter of 3 mm. A bank of low-inductance capacitors varying between 40 and 240 nF charged up to 20 kV through the tube, was used to produce the gas excitation. Below 2000 Å the light radiation was analysed using a 3 m normal incidence vacuum spectrograph with a plate factor in the first diffraction order of 2.77 Å  $\rm mm^{-1}.$  Kodak SWR plates were used to record the spectra and C, N, O and known lines of Xe provided internal wavelength standards. To record the spectra in the UV-visible region, a 3.4 m Ebert plane grating spectrograph at CIOp was used. The grating has 600 lines/mm, corresponding to a plate factor of 5 Å/mm in the first diffraction order. Kodak 103 a-O and Kodak 103 a-F plates were used to record the spectra in the first, second and third diffraction order. Thorium lines from an electrodeless discharge were superimposed on the spectrograms and served as reference lines [20, 21].

To measure the spectrograms a photoelectric semiautomatic Grant comparator was used. We studied the behaviour of the spectral line intensities as a function of pressure and discharge voltage to distinguish between different states of ionisation. In our spectrograms of xenon, as the pressure decreases, higher ionic species tend to appear [16]. In this way, observing the behaviour of the spectral lines and using known lines of Xe VII, we were able to distinguish the different ionic states of our xenon spectra. The accuracy of the wavelength values is estimated to be  $\pm 0.02$  Å in the VUV region. In the UV–visible range the uncertainty is estimated to be  $\pm 0.01$  Å in the first diffraction order.

<sup>&</sup>lt;sup>a</sup> e-mail: gallardom@ciop.unlp.edu.ar

<sup>&</sup>lt;sup>b</sup> e-mail: faustob@ciop.unlp.edu.ar

<sup>&</sup>lt;sup>c</sup> e-mail: monicar@ciop.unlp.edu.ar

<sup>&</sup>lt;sup>d</sup> e-mail: jreyna@ciop.unlp.edu.ar

Table 1. New and adjusted energy values of the  $5s^2$ ,  $5p^2$ , 5s5d, 5s6d, 5s6s, 5s7s and 5s6p, 5s7p, 5p5d, 5s5f configurations of Xe VII.

Desig	nation	$E_{\rm exp}({\rm cm}^{-1})^{\rm a}$	$E_{\rm calc}({\rm cm}^{-1})^{\rm b}$	Percentage composition <sup>c</sup>
$5s^2$	$^{1}S_{0}$	0	0	98
$5p^2$	$^{1}S_{0}$	$278113^{\rm n}$	278073	90 $5p^2$ <sup>3</sup> P + 9 $5p^2$ <sup>1</sup> S
	$^{1}\mathrm{D}_{2}$	236100	236267	$55 \ 5p^2 \ {}^{1}\text{D} + 38 \ 5p^2 \ {}^{3}\text{P} + 6 \ 5s5d \ {}^{1}\text{D}$
5s5d	$^{1}\mathrm{D}_{2}$	307545	307532	83 5s5d $^{1}\text{D} + 10 5p^{2} {}^{1}\text{D}$
5s6d	$^{3}\mathrm{D}_{1}$	475967	475933	100
	$^{3}\mathrm{D}_{2}$	476239	476229	99
	$^{3}\mathrm{D}_{3}$	476784	476827	100
	$^{1}\mathrm{D}_{2}$	$479665^{n}$	479665	96
5s6s	$^{1}S_{0}$	$363511^{\mathrm{n}}$	363511	100
	$^{3}S_{1}$	354833	354833	100
5s7s	$^{1}S_{0}$	$510731^{\rm n}$	510736	100
	$^{3}S_{1}$	506240	506230	100
5s6p	$^{3}P_{0}$	400486	400448	98
	$^{3}P_{1}$	400876	400912	$77 \ 5s6p \ ^{3}P + 21 \ 5s6p \ ^{1}P$
	$^{3}P_{2}$	408088	407973	$69  5s6p  {}^{3}\mathrm{P} + 18  5p5d  {}^{1}\mathrm{D} + 6  5p5d  {}^{3}\mathrm{D}$
	$^{1}P_{1}$	407801	407878	70 $5s6p$ <sup>1</sup> P + 14 $5s6p$ <sup>3</sup> P + 8 $5p5d$ <sup>3</sup> D + 5 $5p5d$ <sup>3</sup> D
5p5d	${}^{3}F_{2}$	$394360^{\rm n}$	394488	$80  5p5d  {}^{3}\mathrm{F} + 17  5p5d  {}^{1}\mathrm{D}$
	${}^{3}F_{3}$	$402505^{\rm n}$	402230	$87 \ 5p5d \ ^{3}\text{F} + 8 \ 5p5d \ ^{3}\text{D}$
	${}^{3}F_{4}$	$414074^{\rm n}$	414005	98
	$^{1}F_{3}$	438427	438837	$78 5p5d {}^{1}F + 13 5s {}^{1}F + 7 5p5d {}^{3}D$
	$^{3}\mathrm{D}_{1}$	$410500^{\rm n}$	410636	$61 \ 5p5d \ ^{3}D + 15 \ 5p5d \ ^{3}P + 9 \ 5s6p^{1}P + 8 \ 5p5d \ ^{1}P + 5s6p \ ^{3}P$
	$^{3}D_{2}$	$418483^{\rm n}$	418287	$37 \ 5p5d^{-1}D + 34 \ 5p5d^{-3}D + 17 \ 5p5d^{-3}P + 10 \ 5p5d^{-3}F$
	$^{3}D_{3}$	423840	423829	$84 \ 5p5d \ ^{3}\text{D} + 11 \ 5p5d \ ^{3}\text{F}$
	$^{1}\mathrm{D}_{2}$		404209	$29 5p5d {}^{3}P + 28 5s6p {}^{3}P + 24 5p5d {}^{1}D + 13 5p5d {}^{3}D + 6 5p5d {}^{3}F$
	$^{3}P_{0}$	$424921^{\rm n}$	424982	98
	$^{3}P_{1}$	425406	425251	$74 \ 5p5d \ ^{3}P + 24 \ 5p5d \ ^{3}D$
	$^{3}P_{2}$	$426105^{\rm n}$	425898	$50 \ 5p5d^{3}P + 45 \ 5p5d^{3}D$
	$^{1}P_{1}$	443192	442955	89 5 $p5d$ <sup>1</sup> P + 6 5 $p5d$ <sup>3</sup> D
5s5f	${}^{3}F_{2}$	$463954^{\rm n}$	463751	99
	${}^{3}F_{3}$	$463810^{\rm n}$	463955	99
	${}^{3}F_{4}$		464227	98
	$^{1}F_{3}$		468322	87
5s7p	$^{3}P_{0}$	527044	526884	100
	${}^{3}P_{1}$	527500	527361	91 $5s7p$ <sup>3</sup> P + 9 $5s7p$ <sup>1</sup> P
	$^{3}P_{2}$	529348	529620	99
	$^{1}P_{1}$	$531766^{\rm n}$	531779	90 $5s7p$ <sup>1</sup> P + 8 $5s7p$ <sup>3</sup> P

<sup>a</sup> Uncertainties of the adjusted experimental energy level values less than  $4 \text{ cm}^{-1}$ .

<sup>b</sup> Calculated energy level values obtained using the fitted energy parameters.

<sup>c</sup> Percentages below 5% have been omitted.

<sup>n</sup> New energy values.

## **3** Analysis

Theoretical predictions of the structure of the configurations were used in the analysis. The predictions were obtained by diagonalizing the energy matrices with appropriately scaled relativistic Hartree-Fock (HFR) values for the energy parameters. The computer code developed by Cowan [22] was used. Comparisons along the Cd I isoelectronic sequence were also used. Table 2 shows the new classified transitions. The intensities of the transitions are visually estimations of plate blackening.

The adjusted experimental energy level values derived from the observed lines are given in Table 1. The level values were determined in an iterative procedure where the wavenumbers of the observed lines are weighted in accordance to their estimated uncertainties. In our case the uncertainties of the adjusted experimental energy level values are generally less than 4 cm<sup>-1</sup>. All level designations

Intoncity	$\lambda$ , $(\hat{\lambda})$ in manufactor	$\sigma_{\rm a}$ ( $m^{-1}$ )	$\sigma \to (cm^{-1})^{b}$	Transition
1 nuensity	Aobs (A) III vacuum	$\frac{\sigma_{\rm obs}}{277}$	$\sigma_{\rm cal}$ (CIII )	5m <sup>2</sup> <sup>3</sup> D <sub>2</sub> F <sub>2</sub> 7m <sup>3</sup> D
1	900.39 400.99	211 008	490	$\partial p$ $r_2 = \partial s (p \cdot r_2)$
2	409.83	244 004	3 994 E2	$5s5a$ $D_1-5s7p$ $P_1$ $5s5a$ $3D$ $5s7m$ $^1D$
1	411.42 /19.05	240 001 241 576	00 E	$5_{250}u D_{2} = 5_{25}tp P_{1}$
1	415.95	241 070	0 79	$5s5a$ $D_1-5stp$ $P_2$ $5n^2$ $^3P_2$ $5n^5f$ $^3P_2$
1	430.10	229 203	12	$5p = r_1 - 5s 5j = r_2$ 5 s 5 d <sup>1</sup> D = 5 s 7m <sup>1</sup> D
1	440.01	224 210	21	$5a5d \ ^{1}D_{2} - 5a7p \ ^{3}P_{1}$
1	450.85	221 805	0	$5s5u D_2 - 5s7p T_2$ $5s5v D_2 - 5s6s T_2$
1 10	404.00	220 230	0	5s5p F1- $5s0s$ 50 $5n^2$ <sup>1</sup> D. $5n5d$ <sup>1</sup> D.
10	402.00	207 091	4	$5p D_2 5p5a T_1$ $5e5n {}^{3}P_{1} - 5e5d {}^{1}D_{2}$
8	404.25	207 091	7	$5n^2 {}^1\text{D}_2 - 5n^5d {}^1\text{F}_2$
0	494.20 522.66	101 320	1 30	$5p^{2} D_{2} - 5p5d^{-1}P_{1}$
1	524.30	191 525	53 94	$5p^{2} {}^{3}P_{1} = 5p5d {}^{3}P_{1}$
-± _/	526 33	190.005	24 89.005	$5p^{2-1}D_{2} - 5n5d^{-3}P_{2}$
-± _/	528.24	189 208	6	$5p^{2-1}D_2 = 5p5d^{-3}P_2$
-+ 7	543 10	184 128	7	$5p^2 {}^{3}P_0 - 5s6n {}^{1}P_1$
1	544 08	183 797	801	$5p^2 {}^{3}P_1 - 5n5d {}^{3}D_2$
2	567 59	176 184	2	$5_{F}$ $^{1}$ $^{1}$ $^{1}$ $^{1}$ $^{1}$ $^{1}$ $^{1}$ $^{1}$ $^{2}$ $^{1}$ $^{2}$ $^{1}$ $^{2}$ $^{1}$ $^{2}$ $^{1}$
4	570.63	175245	- 1	$5s5d {}^{3}D_{2}-5s5f {}^{3}F_{2}$
3	571.10	175 101	097	$5s5d {}^{3}D_{2}-5s5f {}^{3}F_{2}$
1	573.38	174404	0	$5n^2 {}^1D_2 - 5n5d {}^3D_1$
1	573.86	174259	2	$5p^2 {}^{3}P_2 - 5n5d {}^{3}P_2$
5	574.38	174 101	098	$5s5p {}^{3}P_{2}-5s5d {}^{3}D_{1}$
4	576.48	173467	71	$5s5d {}^{3}D_{3}-5s5f {}^{3}F_{3}$
8	577.63	173121	19	$5p^2 {}^3P_1 - 5s6p {}^1P_1$
3	579.15	172667	7	$5s6s {}^{3}S_{1} - 5s7p {}^{3}P_{1}$
4	580.66	172 218	1	$5s6s {}^{3}S_{1} - 5s7p {}^{3}P_{0}$
7	694.00	160 406	5	$5p^{2-1}D_2-5p5d^{-3}F_3$
2	601.71	166193	4	$5p^2 {}^3P_1 - 5s6p {}^3P_1$
3	605.79	165074	9	$5p^2 {}^1S_0 - 5p5d {}^1P_1$
2	639.37	156403	9	$5s5d \ ^{1}D_{2}-5s5f \ ^{3}F_{2}$
4	639.94	156265	5	$5s5d {}^{1}D_{2}-5s5f {}^{3}F_{3}$
4	640.03	156243	35	$5p^2 {}^{3}P_2 - 5s6p {}^{3}P_2$
1	671.02	149027	3	$5p^2 {}^3P_2 - 5s6p {}^3P_1$
4	675.28	148 087	8	$5s5d {}^{3}D_{3}-5p5d {}^{1}F_{3}$
5	687.53	145448	52	$5s5p {}^{1}P_{1} - 5s5d {}^{3}D_{2}$
5	691.98	144513	1	$5s5p {}^{1}P_{1} - 5s5d {}^{3}D_{1}$
4	701.74	142503	7	$5p^2 {}^3P_2 - 5p5d {}^3F_2$
1	727.85	137391	2	$5s5d {}^{3}D_{2}-5p5d {}^{3}P_{2}$
1	729.11	137154	49	$5s5d {}^{3}D_{1} - 5p5d {}^{3}P_{0}$
5	731.56	136694	4	$5s5d {}^{3}D_{2}-5p5d {}^{3}P_{1}$
8	737.21	135647	7	$5s5d {}^{1}D_{2} - 5p5d {}^{1}P_{1}$
8	737.21	135647	9	$5s5p {}^{3}P_{1} - 5p^{2} {}^{1}D_{2}$
3	741.56	134851	2	$5s5p {}^{1}P_{1} - 5p^{2} {}^{1}S_{0}$
8	749.03	133506	1	$5s5d {}^{3}D_{3}-5p5d {}^{3}D_{3}$
7	764.02	130 887	2	$5s5d {}^{1}D_{2}-5p5d {}^{1}F_{3}$
4	765.02	130716	1	$5s5d \ {}^{3}D_{1} - 5p5d \ {}^{3}D_{2}$
5	770.62	129766	70	$5s5d \ {}^{3}D_{2}-5p5d \ {}^{3}D_{2}$
5	771.07	129690	88	$5p^2 {}^1S_0 - 5s6p {}^1P_1$
7	780.34	128 149	4	$5s5d {}^{3}D_{3}-5p5d {}^{3}D_{2}$

 Table 2. Classified lines of Xe VII.

Intensity	$\lambda_{\rm obs}$ (Å) in vacuum	$\sigma_{\rm obs}~({\rm cm}^{-1})$	$\sigma_{\rm cal} \ ({\rm cm}^{-1})^{\rm b}$	Transition
3	808.18	123 735	5	$5s5d \ {}^{3}D_{3}-5p5d \ {}^{3}F_{4}$
3	814.81	122728	8	$5s5d {}^{3}D_{1}$ - $5p5d {}^{3}D_{1}$
4	831.16	120314	6	$5s5d {}^{3}D_{1}$ - $5s6p {}^{3}P_{2}$
5	837.20	119375	6	$5s5d {}^{3}D_{2}-5s6p {}^{3}P_{2}$
4	839.75	119083	8	$5s5d {}^{3}D_{2}-5s6p {}^{1}P_{1}$
2	848.43	117865	1	$5s5d {}^{1}D_{2}-5p5d {}^{3}P_{1}$
6	849.26	117750	49	$5s5d {}^{3}D_{3}-5s6p {}^{3}P_{2}$
4	874.69	114326	34	$5s5d \ {}^{3}D_{2}-5p5d \ {}^{1}D_{2}$
6	884.13	113106	4	$5s5d {}^{3}D_{1}-5s6p {}^{3}P_{1}$
6	887.20	112714	4	$5s5d {}^{3}D_{1}-5s6p {}^{3}P_{0}$
5	891.55	112164	6	$5s5d \ {}^{3}D_{3}-5p5d \ {}^{3}F_{3}$
5	891.55	112164	4	$5s5d {}^{3}D_{2}-5s6p {}^{3}P_{1}$
3	938.23	106584	8	$5s5d \ {}^{3}D_{1}$ - $5p5d \ {}^{3}F_{2}$
2	946.49	105654	47	$5s5d {}^{3}D_{2}-5p5d {}^{3}F_{2}$
2	971.52	102932	0	$5s6p {}^{1}P_{1} - 5s7s {}^{1}S_{0}$
3	1044.54	95736	40	$5p5d$ $^{3}D_{1}-5s7s$ $^{3}S_{1}$
6	1213.29	82421	4	$5p5d$ ${}^{3}F_{2}$ - $5s6d$ ${}^{3}D_{3}$
4	1225.46	81602	7	$5p5d$ ${}^{3}F_{2}$ - $5s6d$ ${}^{3}D_{1}$
4	1229.83	81312	9	$5p5d$ $^{3}P_{0}$ - $5s7s$ $^{3}S_{1}$
7	1243.58	80413	3	$5s5p {}^{1}P_{1} - 5p^{2} {}^{3}P_{0}$
2	1247.90	80135	5	$5p5d {}^{3}P_{2}-5s7s {}^{3}S_{1}$
2	1254.92	79686	1	$5s6s {}^{1}S_{0} - 5p5d {}^{1}P_{1}$
3	1296.05	77158	60	$5p5d$ ${}^{3}F_{3}$ - $5s6d$ ${}^{1}D_{2}$
2	1326.93	75362	3	$5s6p \ {}^{3}P_{1}$ – $5s6d \ {}^{3}D_{2}$
6	1331.63	75096	1	$5s6p \ {}^{3}P_{1} - 5s6d \ {}^{3}D_{1}$
2	1356.15	73738	4	$5p5d$ $^3\mathrm{F}_35s6d$ $^3\mathrm{D}_2$
4	1391.49	71865	4	$5s6p \ ^{1}P_{1} - 5s6d \ ^{1}D_{2}$
2	1396.96	71584	77	$5s6p \ ^{3}P_{2}$ - $5s6d \ ^{1}D_{2}$
4	1426.72	70091	88	$5s6s~^{3}S_{1}$ – $5p5d~^{3}P_{0}$
2	1445.81	69165	5	$5p5d~^{3}D_{1}$ – $5s6d~^{1}D_{2}$
3	1455.68	68696	6	$5s6p \ {}^{3}P_{2} - 5s6d \ {}^{3}D_{3}$
2	1460.95	68449	38	$5s6p~^{1}P_{1}$ – $5s6d~^{3}D_{2}$
4	1473.25	67877	9	$5s6p \ {}^{3}P_{2}$ – $5s6d \ {}^{3}D_{1}$
2	1480.57	67542	39	$5p5d~^{1}P_{1}$ - $5s7s~^{1}S_{0}$
8	1527.40	65471	67	$5p5d$ $^{3}D_{1}$ – $5s6s$ $^{3}D_{1}$
6	1586.00	63052	48	$5p5d~^1\mathrm{P}_15s7s~^3\mathrm{S}_1$
10	1594.63	62711	0	$5p5d~^{3}F_{4}$ – $5s6d~^{3}D_{3}$
2	1615.65	61895	5	$5s6s {}^{1}S_{0}$ - $5p5d {}^{3}P_{1}$
8	1715.23	58301	1	$5p5d\ {}^{3}\mathrm{D}_{2}5s6d\ {}^{3}\mathrm{D}_{3}$
8	1731.39	57757	8	$5p5d\ {}^{3}\mathrm{D}_{2}5s6d\ {}^{3}\mathrm{D}_{2}$
7	1791.17	55829	5	$5p5d~^{3}D_{3}$ – $5s6d~^{1}D_{2}$
8	1843.20	54254	9	$5p5d \ {}^{3}P_{1}-5s6d \ {}^{1}D_{2}$
7	1867.08	53560	0	$5p5d\ {}^{3}P_{2}$ – $5s6d\ {}^{1}D_{2}$
7	1888.81	52943	4	$5p5d\ {}^{3}\mathrm{D}_{3}5s6d\ {}^{3}\mathrm{D}_{3}$
4	1902.47	52563	4	$5s6d \ {}^{3}D_{3}$ – $5s7p \ {}^{3}P_{2}$
6	1919.19	52105	1	$5s6d \ ^{1}D_{2}-5s7p \ ^{1}P_{1}$
7	1950.78	51262	1	$5s6d {}^{3}D_{2}-5s7p {}^{3}P_{1}$
7	1957.79	51078	7	$5s6d {}^{3}D_{1}-5s7p {}^{3}P_{0}$
7	1958.87	51050	46	$5p5d~^3\mathrm{P}_05s6d~^3\mathrm{D}_1$
7	1967.07	50837	3	$5p5d \ {}^{3}P_{1}$ – $5s6d \ {}^{3}D_{2}$

 Table 2. Continued.

Intensity	$\lambda_{\rm obs}$ (Å) in vacuum	$\sigma_{ m obs}~( m cm^{-1})$	$\sigma_{\rm cal} \ ({\rm cm}^{-1})^{\rm b}$	Transition	
5	1973.41	50674	9	$5p5d \ {}^{3}P_{2}$ – $5s6d \ {}^{3}D_{3}$	
3	1994.59	50136	9	$5p5d~^3\mathrm{P}_25s6d~^3\mathrm{D}_2$	
8	2090.27	47841	35	$5s6d \ ^{1}D_{2}$ - $5s7p \ ^{3}P_{1}$	
2	2128.01	46992	89	$5s6s~^{1}S_{0}$ - $5p5d~^{3}D_{1}$	
	$\lambda_{ m air}$				
5	2257.13	44290.5	0.0	$5s6s {}^{1}S_{0} - 5s6p {}^{1}P_{1}$	
9	3916.42	25526.3	6.0	$5s7s$ ${}^{3}S_{1}$ - $5s7p$ ${}^{1}P_{1}$	
4	4326.35	23107.7	8.0	$5s7s \ {}^{3}S_{1} - 5s7p \ {}^{3}P_{2}$	
2	4702.32	21260.2	0.0	$5s7s~^3\mathrm{S}_15s7p~^3\mathrm{P}_1$	
8	4752.68	21034.9	5.0	$5s7s$ $^{1}S_{0}$ - $5s7p$ $^{1}P_{1}$	
12	4805.44	20804.0	4.0	$5s7p {}^{3}P_{0}-5s7s {}^{3}S_{1}$	

Table 2. Continued.

Intensity, observed lines intensities from visual estimation.

<sup>b</sup> Calculated wavenumber. Only the last digits which differ from the observed ones are given.

in Table 1 are in LS notation, and in the same table we present the percentage composition of the levels.

In the even parity configurations the  $5p^2$  <sup>1</sup>S<sub>0</sub> energy level value reported by Larsson et al. [8] was rejected, and we propose a new energy level in  $278\,113$  cm<sup>-1</sup>, supported by three new transitions with the 5s5p  ${}^{1}P_{1}$ , 5s6p  ${}^{1}P_{1}$ , and the 5p5d  ${}^{1}P_{1}$  energy levels respectively. The new 5s5d <sup>1</sup>D<sub>2</sub> energy level value proposed by Larsson *et al.* [8] was confirmed and adjusted by adding seven new transitions shown in Table 1. We reject the 5s6s  ${}^{1}S_{0}$  energy level value reported by these same authors and propose a new value in  $363511 \text{ cm}^{-1}$ , supported by five new transitions with the 5s5p  ${}^{1}P_{1}$ , 5s6p  ${}^{1}P_{1}$  and the 5p5d  ${}^{1}P_{1}$ ,  ${}^{3}P_{1}$ and  ${}^{3}D_{1}$  energy levels respectively. We establish a new 5s6d <sup>1</sup>D<sub>2</sub> energy level value in 479665 cm<sup>-1</sup>, supported by nine new transitions. We also establish a new 5s7s  ${}^{1}S_{0}$ energy level value in  $510731 \text{ cm}^{-1}$ , supported by three new transitions.

In the odd parity configurations, all the previously known energy level values for the 5s5p, 5s6p, and 5s7pconfigurations were confirmed, and we propose a new  $5s7p^{-1}P_1$  energy level value in 531766 cm<sup>-1</sup> supported by six new transitions with the  $5s5d {}^{3}D_{1}$ ,  ${}^{3}D_{2}$ ,  $5s5d {}^{1}D_{2}$ ,  $556d^{-1}D_2$ , and  $5s7s^{-1}S_0$ ,  ${}^{3}S_1$  energy levels. We establish two new  $5s5f^{-3}F_2$  and  $5s5f^{-3}F_3$  energy level values in  $463\,954$  cm<sup>-1</sup> and  $463\,810$  cm<sup>-1</sup> respectively, supported by the transitions shown in Table 1. We were not able to find transitions that permit to establish the 5s5f  ${}^{3}F_{4}$  and 5s5f  ${}^{1}F_{3}$  energy level values. We rejected the 5p5d  ${}^{3}P_{2}$ energy level value reported in the paper of reference [7], and propose a new value in  $426\,105 \text{ cm}^{-1}$ , supported by seven transitions shown in Table 1. We also establish a new 5p5d <sup>3</sup>P<sub>0</sub> energy level value in 424921 cm<sup>-1</sup>, supported by four transitions shown in Table 1. The energy level values for the 5p5d  $^{1}P_{1}$ ,  $^{1}F_{3}$ ,  $^{3}P_{1}$ , and  $^{3}D_{3}$  reported in the paper of reference [7] were confirmed.

We reject the 5p5d  ${}^{3}D_{1}$  and  ${}^{3}D_{2}$  energy level values reported in the work of reference [7], and propose new values in 410 500 cm<sup>-1</sup> and 418 483 cm<sup>-1</sup> respectively. These values are supported by six new transitions each other shown

in Table 1. We also reject the energy level values reported in reference [7] for all the J values belonging to the  $5p5d \,{}^3F$ multiplet, and propose a new value in 394 360 cm<sup>-1</sup> for the  ${}^3F_2$ , a new value in 402 505 cm<sup>-1</sup> for the  ${}^3F_3$ , supported by five new transitions each other; and a new value in 414074 cm<sup>-1</sup> for the  ${}^3F_4$ , supported by two transitions with the  $5s5d \,{}^3D_3$  and the  $5s6d \,{}^3D_3$  energy levels.

### **4** Discussion

When interpreting the observed energy-level structures we performed least-squares fits between the experimental energy-level values and the eigenvalues obtained when diagonalizing the energy matrices.

To study the even parity configuration theoretically we have considered preliminary the complete  $5s^2 + 5p^2 +$  $5s(6s + 7s + 5d + 6d) + 5p(4f + 5f) + 5p6p + 4f^2 + 5d^2 + 5d^2$  $4d^95s^2(6s+5d)$  energy matrix, and we observed that the more high configurations and also the core excited ones interact with the configurations here analyzed. The results of the parametric calculations for the even configurations are presented in Table 3. We included in the fitting the 4f5p configuration in accordance with the procedure reported calculations along the isoelectronic sequence [13]. The scaled Hartree-Fock factor was 0.85 for all the parameters, except for  $\zeta_{nl}$  and for  $E_{av}$  that was 0.95 and 1.00 respectively. In our case all of the experimental level values of the 4f5p configuration are unknown, and due to this, all parameters of this configuration were held fixed in the least-squares fit calculation. Besides, all the configuration interaction integrals were held fixed in the calculation, at 0.85 of their HF values, except for the  $5s^2-5p^2$ .  $5p^2 - 5s5d$ ,  $5p^2 - 5s6d$ , 5s5d - 5s6d, and 5s6d - 4f5p, that were held fixed in the calculation at 0.70 of their HF values, in order to obtain a better parametric fit.

Ab initio calculations to study the odd parity configurations considering the  $5s(5p + 6p + 7p) + 5s(4f + 5f) + 5p(5d+6d) + 5p6s + 5d6p + 5d(4f + 5f) + 4d^95s^25p$  energy matrices were carried out. The results of the parametric

Configuration	Parameter	HF value	Fitted value	$\mathrm{F/HF}(^*)$
$5s^{2}$	$E_{av}$	00	3667	
$5p^2$	$E_{av}$	238723	$246066\pm59$	$1.030\pm0.001$
	$\zeta_{5p}$	10573	$11417\pm64$	$1.079\pm0.006$
	$F^2(5p,5p)$	59133	$48322\pm318$	$0.817 \pm 0.005$
5s5d	$E_{av}$	283684	$293692\pm64$	$1.035\pm0.001$
	$\zeta_{5d}$	849	$1062\pm65$	$1.250\pm0.076$
	$G^{2}(5s, 5d)$	38852	$35253\pm402$	$0.907 \pm 0.010$
5s6d	$E_{av}$	467624	$476872\pm59$	$1.019\pm0.001$
	$\zeta_{6d}$	381	362 (FIX)	0.950
	$G^2(5s, 6d)$	9150	$4915\pm348$	$0.537 \pm 0.038$
5s6s	$E_{av}$	347141	$357020\pm92$	$1.028\pm0.001$
	$G^0(5s,6s)$	5220	$4353\pm82$	$0.833 \pm 0.015$
5s7s	$E_{av}$	494171	$507326 \pm 92$	$1.026\pm0.001$
	$G^0(5s,7s)$	1792	$2193\pm82$	$1.223\pm0.045$
4f5p	$E_{av}$	388137	388137~(FIX)	1.000
	$\zeta_{4f}$	254	242 (FIX)	0.950
	$\zeta_{5p}$	10165	9657 (FIX)	0.950
	$F^2(4f,5p)$	51236	$43551~({\rm FIX})$	0.850
	$G^2(4f, 5p)$	35088	29825~(FIX)	0.850
	$G^4(4f, 5p)$	26040	22134~(FIX)	0.850
Configuration i	nteraction integrals			
$5s^2 - 5p^2$	$R^1(5s5s, 5p5p)$	76811	$53767~(\mathrm{FIX})$	0.700
$5s^2 - 5s6s$	$R^0(5s5s, 5s6s)$	3890	3307 (FIX)	0.850
$5s^2 - 5s7s$	$R^0(5s5s,5s7s)$	1899	1614 (FIX)	0.850
$5p^2 - 5s5d$	$R^1(5p5p, 5s5d)$	64976	45483 (FIX)	0.700
$5p^2 - 5s6d$	$R^1(5p5p, 5s6d)$	25549	17884 (FIX)	0.700
$5p^2 - 5s6s$	$R^1(5p5p, 5s6s)$	-318	-271 (FIX)	0.850
$5p^2 - 5s7s$	$R^1(5p5p, 5s7s)$	-1862	-1583 (FIX)	0.850
$5p^2 - 4f5p$	$R^2(5p5p, 4f5p)$	-42609	-36218 (FIX)	0.850
$5s5d{-}5s6d$	$R^0(5s5d, 5s6d)$	0	0 (FIX)	
	$R^2(5s5d, 6d5s)$	17211	12048~(FIX)	0.700
$5s5d{-}4f5p$	$R^3(5s5d, 4f5p)$	-30636	-26041 (FIX)	0.850
	$R^1(5s5d, 5p4f)$	-37604	-31964 (FIX)	0.850
$5s6d{-}4f5p$	$R^3(5s6d, 4f5p)$	-13701	-9590 (FIX)	0.700
	$R^1(5s6d, 5p4f)$	-14362	-10053 (FIX)	0.700
5s7s-4f5p	$R^0(5s7s, 4f5p)$	0	0 (FIX)	
	$R^{0}(5s7s, 5p4f)$	2990	2542 (FIX)	0.850

Table 3. Energy parameters  $(cm^{-1})$  for the studied even parity configurations of XeVII.

(\*) Fitted value/Hartree-Fock value.

The standard deviation of the fit is  $117 \text{ cm}^{-1}$ .

calculations are showed in Table 4, where the 5s7p configuration and the rest of the fitted configurations reported in reference [13] are included. In our calculations the radial parameters belonging to the 4f5s and the  $4d^95s^25p$ configuration were held fixed at 0.85, 0.95 and 1.00 for the  $G^k$ ,  $\zeta_{nl}$  and  $E_{av}$  respectively. In the 5s6p configuration the  $G^1(5s, 6p)$  integral was optimized and fixed at 0.25 of its HF value. This value is in good agreement with the extrapolated value obtained along the isoelectronic sequence, reported in reference [9]. All the configuration interaction integrals were fixed at 0.85 of their HF values, except for the 5s5p-5p5d, and 5p5d-5s5f integrals, fixed at 0.70 and 0.75 respectively.

Configuration	Paramotor	HF value	Fitted value	F/HF
5 a5m	F	108 320	$\frac{117703 \pm 153}{117703 \pm 153}$	1000000000000000000000000000000000000
0s0p	$L_{av}$	106520	$117705 \pm 155$ $11681 \pm 944$	$1.030 \pm 0.001$ $1.006 \pm 0.022$
	$C^{5p}$	76 892	$11001 \pm 244$	$1.090 \pm 0.022$ 0.726 \pm 0.007
E of m	$G_{(3s, 3p)}$	10823	$30300 \pm 373$	$0.730 \pm 0.007$ 1.022 $\pm 0.001$
5s0p	$L_{av}$	390 090	$403461 \pm 192$ $4006 \pm 220$	$1.025 \pm 0.001$
	$\zeta_{6p}$	0912 0505	$4290 \pm 250$ 014C (EIX)	$1.061 \pm 0.000$
5 o 7 m	G(3s, 0p)	8080 E19497	2140 (FIA)	0.230 1.021 $\pm$ 0.001
5 <i>s i p</i>	$L_{av}$	010407 1087	$329300 \pm 141$ 1900 (FIV)	$1.021 \pm 0.001$
	$\zeta_{7p}$	1967	1000 (FIA)	0.900
F F 1	$G^{2}(5s, tp)$	31/7	$4122 \pm 551$	$1.297 \pm 0.173$
$_{\mathrm{D}p\mathfrak{I}a}$	$E_{av}$	400 352	$417807 \pm 103$ 10.277 $\pm 0.01$	$1.028 \pm 0.001$
	$\zeta_{5p}$	10739	$12377 \pm 201$	$1.150 \pm 0.024$ 1.100 $\pm 0.100$
	$\zeta_{5d}$ $E^2(\Gamma_{11},\Gamma_{1})$	800	$970 \pm 173$	$1.120 \pm 0.199$
	F(3p, 3a)	49801	$32000\pm970$	$0.052 \pm 0.019$
	$G^{2}(5p, 5d)$	58 636	$37310 \pm 495$	$0.636 \pm 0.008$
	$G^{o}(5p, 5d)$	37 484	$22037\pm1132$	$0.587 \pm 0.030$
4f5s	$E_{av}$	268 590	268590 (F1X)	1.000
	$\zeta_{4f}$	248	235 (FIX)	0.950
	$G^{o}(4f, 5s)$	36 157	30733(F1X)	0.850
5s5f	$E_{av}$	452 417	$463602\pm199$	$1.024 \pm 0.001$
	$\zeta_{5f}$	89	85 (FIX)	0.950
4 19 - 2 -	$G^{\circ}(5s,5f)$	4471	3800 (FIX)	0.850
$4d^{s}5s^{2}5p$	$E_{av}$	532110	540110 (FIX)	1.015
	$\zeta_{4d}$	6274	5960 (FIX)	0.950
	$\zeta_{5p}$	11505	10930~(F1X)	0.950
	$F^{(2)}(5p, 5d)$	41955	35662 (F1X)	0.850
	$G^{1}(5p, 5d)$	12587	10699 (FIX)	0.850
~ ~	$G^{3}(5p, 5d)$	12281	10438~(FIX)	0.850
Configuration in	teraction integrals			
5s5p-5s6p	$R^{0}(5s5p, 5s6p)$	3246	2759 (F1X)	0.850
	$R^{1}(5s5p, 6p5s)$	19080	16218~(FIX)	0.850
5s5p-5s7p	$R^{0}(5s5p, 5s7p)$	1644	1397 (F1X)	0.850
_	$R^{1}(5s5p, 7p5s)$	9589	8151 (FIX)	0.850
5s5p-5p5d	$R^{1}(5s5p, 5p5d)$	65 650	45955~(FIX)	0.700
-0.2	$R^{2}(5s5p, 5d5p)$	48308	41062~(F1X)	0.850
$5s5p-4d^{s}5s^{2}5p$	$R^{2}(4d4d, 4d5s)$	-8155	-6931(FIX)	0.850
	$R^{2}(4d5p, 5s5p)$	-22705	-19299 (FIX)	0.850
	$R^{1}(4d5p, 5p5s)$	-20004	-17004 (FIX)	0.850
5s6p-5s7p	$R^{0}(5s6p, 5s6p)$	0	0 (FIX)	
	$R^{1}(5s6p, 6p5s)$	5095	4331 (FIX)	0.850
5s6p-5p5d	$R^{1}(5s6p, 5p5d)$	-3559	-3025 (FIX)	0.850
	$R^{2}(5s6p, 5d5p)$	6469	5498 (FIX)	0.850
$5s6p - 4d^95s^25p$	$R^{2}(4d6p, 5s5p)$	-5936	-5046 (FIX)	0.850
	$R^{1}(4d6p, 5s5p)$	-4762	-4047 (FIX)	0.850
5s7p-5p5d	$R^{1}(5s7p, 5p5d)$	-3106	-2640 (FIX)	0.850
	$R^{2}(5s7p, 5p5d)$	1896	1611 (FIX)	0.850
$5s7p - 4d^95s^25p$	$R^{2}(4d7p, 5s5p)$	-2913	-2476 (FIX)	0.850
	$R^{1}(4d7p, 5p5s)$	-2221	-1888 (FIX)	0.850
5p5d - 4f5s	$R^2(5p5d, 4f5s)$	-35170	-29894 (FIX)	0.850
* •	$R^{1}(5p5d, 5s4f)$	-39041	-33185 (FIX)	0.850
5p5d-5s5f	$R^{1}(5p5d, 5s5f)$	29021	21 766 (FIX)	0.750
÷ v	$R^2(5p5d, 5f5s)$	10765	9150 (FIX)	0.850
$5p5d - 4d^95s^25n$	$R^{2}(4d5d, 5s5s)$	-16644	-14148 (FIX)	0.850
4f5s - 5s5f	$R^{3}(4f5s.5s5f)$	-125	-106 (FIX)	0.850
J - J	$R^{0}(4f5s.5f5s)$	0	0 (FIX)	
$4f5s - 4d^95s^25n$	$R^{2}(4d4f, 5s5n)$	23511	19984 (FIX)	0.850
,	$R^{3}(4d4f, 5n5s)$	19363	16 458 (FIX)	0.850
$5s5f - 4d^95s^25n$	$R^{2}(4d5f, 5s5n)$	454	386 (FIX)	0.850
,p	$R^{3}(4d5f, 5p5s)$	2460	2091 (FIX)	0.850

Table 4. Energy parameters  $(cm^{-1})$  for the studied odd parity configurations of XeVII.

The standard deviation of the fit is  $279 \text{ cm}^{-1}$ .

We are deeply grateful to support from Conselho Nacional de Desenvolvimento Científico e Tecnológico, CNPq, Brasil; Fundação de Amparo à Pesquisa do Estado de São Paulo (Fapesp), Brasil; Comisión de Investigaciones Científicas de la Provincia de Buenos Aires (CIC), Argentina, where J.G.R.A., F.B. and M.R. belong as researchers; and Consejo Nacional de Investigaciones Científicas y Técnicas (CONICET), Argentina, is gratefully acknowledged.

#### References

- B.C. Fawcett, B.B. Jones, R. Wilson, Proc. Phys. Soc. Lond. 78, 1223 (1961).
- E.J. Knystautas, J. Sugar, J.R. Roberts, J. Opt. Soc. Am. 69, 1726 (1979).
- 3. R. Hallin et al., Nucl. Instrum. Meth. 202, 41 (1982).
- 4. G. O'Sullivan, J. Phys. B **15**, L765 (1982).
- 5. J. Blackburn et al., J. Opt. Soc. Am. 73, 1325 (1983).
- 6. V. Kaufman, S. Sugar, J. Opt. Soc. Am. B 4, 1919 (1987).
- 7. G.H. Cavalcanti et al., J. Opt. Soc. Am. B 14, 2459 (1997).

- 8. M.O. Larsson et al., Phys. Scripta 51, 69 (1995).
- 9. M. Wang et al., J. Opt. Soc. Am. B 14, 3277 (1997).
- A. Tauheed, Y.N. Joshi, A.F. Zafaran, Phys. Scripta 62, 316 (2000).
- A. Tauheed, Y.N. Joshi, E.H. Pinnington, Phys. Scripta 56, 289 (1997).
- R. Gayasov, Y.N. Joshi, J. Opt. Soc. Am. B 16, 1280 (1999).
- 13. S.S. Churilov, Y.N. Joshi, Phys. Scripta 62, 282 (2000).
- R. Gayasov, Y.N. Joshi, A.N. Ryabtsev, Phys. Scripta 59, 419 (1999).
- 15. J.G. Reyna Almandos et al., Phys. Rev. A 43, 6098 (1991).
- 16. M. Gallardo *et al.*, Phys. Scripta **51**, 737 (1995).
- M. Gallardo *et al.*, J. Quant. Spect. Rad. Transf. **61**, 319 (1999).
- 18. R. Sarmiento et al., J. Phys. B 32, 2853 (1999).
- 19. M. Gallardo et al., Appl. Opt. 28, 4513 (1989).
- 20. F.P.J. Valero, J. Opt. Soc. Am. 58, 484 (1968).
- D. Goorwitch, F.P.J. Valero, A.L. Clua, J. Opt. Soc Am. 59, 971 (1969).
- 22. R.D. Cowan, *The Theory of Atomic Structure and Spectra* (University of California Press, Berkeley, CA, 1981).